Meteorite source regions as revealed by the near-Earth object population

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Spectroscopic and taxonomic information is now available for 1000 near-Earth objects, having been obtained through both targeted surveys (e.g. [1–3]) or resulting from all-sky surveys (e.g. [4]). We first evaluate these results within the framework of taxonomic types in the Bus-DeMeo system [5,6] and subsequently examine meteorite correlations based on spectral and mineralogical analysis (e.g. [7,8]). We correlate our spectral findings with the source region probabilities calculated using the methods of Bottke et al. [9]. The source regions evaluated are Mars Crossers,  $\nu_6$  resonance, 3:1 resonance, the Outer Belt, and Jupiter Family Comets.

In terms of taxonomy, very clear sources are indicated: Q-, Sq-, and S-types most strongly associated with ordinary chondrite meteorites show clear source signatures through the innermost main-belt regions. V-types are relatively equally balanced between  $\nu_6$  and 3:1 resonance sources, consistent with the orbital dispersion of the Vesta family. Asteroid taxonomy classes interpreted as analogous to meteorites with primitive compositions, B- and C-types, show distinct source region preferences for the outer belt and for Jupiter family comets. Most strongly indicated is a Jupiter family comet source for the D-type near-Earth objects, implying a pronounced likelihood that these "asteroidal" bodies are extinct or dormant comets [10]. Similarly, near-Earth objects falling in the spectrally featureless "X-type" category also show a strong outer belt and Jupiter family comet source region preference; even though they lack albedo measurements, they may be interpreted as originating from among "P-type" primitive objects common in the outer belt. Finally the Xe-class of near-Earth objects, which most closely match the spectral properties of enstatite achondrite (aubrite) meteorites, show a source region preference consistent with a Hungaria origin (confirming [11]) by entering near-Earth space through the Mars crossing and  $\nu_6$  resonance pathways.

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References: [1] Lazzarin, M. et al. (2004), Mem. S. A. It. Suppl. 5, 21. [2] Thomas, C. A. et al. (2014), Icarus 228, 217. [3] http://smass.mit.edu/minus.html [4] Hasselmann, P. H., Carvano, J. M., Lazzaro, D. (2011) NASA PDS, EAR-A-I0035-5-SDSSTAX-V1.0. [5] Bus, S.J., Binzel, R.P. (2002). Icarus 158, 146. [6] DeMeo, F.E. et al. (2009), Icarus 202, 160. [7] Dunn et al. (2010) Icarus 208, 789. [8] Dunn et al. (2013) Icarus 222, 273. [9] Bottke, W.F. et al. (2002), Icarus 156, 399. [10] DeMeo, F., Binzel, R. P. (2007) Icarus 194, 436. [11] Gaffey, M. J. et al. (1992) Icarus 100, 95.