

Organic samples produced by ion bombardment of ices for the EXPOSE-R2 mission on the International Space Station

G. Baratta¹, D. Chaput², H. Cottin³, L. Fernandez Cascales¹, M. Palumbo¹, and G. Strazzulla¹

¹INAF-Osservatorio Astrofisico di Catania, via Santa Sofia 78, 95123 Catania, Italy

²Centre National d'Etudes Spatiales (CNES), Centre Spatial de Toulouse, 18 Av. Edouard Belin, 31401 Toulouse Cedex 9, France

³Laboratoire Interuniversitaire des Systemes Atmospheriques (LISA), UMR CNRS 7583, Université Paris Est Creteil et Université Paris Diderot, Institut Pierre Simon Laplace, France

We describe the preparation and characterization (by UV-Vis-IR spectroscopy) of a set of organic samples, stable at room temperature and above, that are part of the experiment “Photochemistry on the Space Station (PSS)” planned to be enclosed in the EXPOSE-R2 mission, which will be conducted on the EXPOSE-R facility, outside the International Space Station (ISS). The organic materials are prepared in the Catania laboratory after 200 keV He⁺ irradiation of icy mixtures, namely N₂:CH₄:CO deposited at 16 K on MgF₂ windows furnished by European Space Agency. It is widely accepted that such kind of materials produced by energetic processing are representative of organic material in astrophysical environments such as, e.g., comets. Once expelled from comets these materials are exposed to solar radiation during their interplanetary journey before they eventually land on the Earth and other planetary objects where they might give a contribution to the chemical and pre-biogenic evolution. In particular our residues contain different chemical groups, including triple CN bonds that are considered relevant to pre-biogenic chemistry (e.g. Palumbo et al., 2000). Therefore the samples will be exposed, for several months, to the solar ultraviolet photons that are a major source of energy to initiate chemical evolution in the Solar System. This will allow analysis of their destruction and evaluation of their lifetime in the interplanetary medium. The samples have three different thicknesses (about 200, 130, 65 nm) that will allow the estimation of the depth profile of destruction (e.g., Baratta et al., 2002). This experiment overcomes the limits of ground tests which do not reproduce exactly the space parameters.

Acknowledgements: This research has been financially supported by the Italian Space Agency contract: PSS (Photochemistry on the Space Station).

References: Baratta G.A., Leto G., Palumbo M.E.: A comparison of ion irradiation and UV photolysis of CH₄ and CH₃OH, *A&A*, 384, 343, 2002; Palumbo, M.E., Strazzulla, G., Pendleton, Y.J., Tielens, A.G.G.M. ROCN species produced by ion irradiation of ice mixtures: comparison with astronomical observations. *ApJ* 534, 801, 2000.