Observations of ammonia in comets with Herschel

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Ammonia is the most abundant nitrogen bearing species in comets. However, it has been scarcely observed in comets due to the weakness of the lines observable from the ground at infrared and centimetre wavelengths. Nevertheless, its main photodissociation product NH₂ has been observed in several comets in the visible. The fundamental rotational $J_K = (1_0 - 0_0)$ transition of NH₃ at 572.5 GHz has been observed in comets since 2004, with the Odin satellite (Biver et al. 2007). In the frame of the Herschel guaranteed time key program "HssO" (Hartogh et al. 2009), ammonia was detected with the HIFI instrument in comets 10P/Tempel 2 (Biver et al. 2012), 45P/Honda- Mrkos-Pajdusakova, 103P/Hartley 2, and C/2009 P1 (Garradd). The hyperfine structure of the line is resolved. We have built a complete excitation model to interpret these observations, including the radial distribution in comet 103P. The derived abundances relative to water are on the order of 0.5 %, similar to the values inferred from visible observations of NH₂.

References: Biver, N., Bockelée-Morvan, D., Crovisier, J., et al. 2007, P&SS, 55, 1058; Biver, N., Crovisier, J., Bockelée-Morvan, D., et al. 2012, A&A, 539, A68; Hartogh, P., Lellouch, E., Crovisier, J., et al. 2009, P&SS 57, 1596.