Thermal emission photometry of three near-Earth asteroids in L' and M'

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We acquired JKL'M' (1.2–4.7 microns; [1]) spectrophotometry of three sub-km near-Earth asteroid radar targets during 2013. Thermal emission measurements in L' and M' were taken in order to characterize the sizes and thermal properties of the targets. Reflectance data in J and K were taken to aid in separating the thermal from the reflected flux in L'. In addition, we used the J band as a reference filter to characterize the rotational variability in the asteroid lightcurves during the thermal observations.

Asteroid (163249) 2002 GT was observed using NIRI on Gemini North [2]. Numerous coordinating observations of this asteroid were made as part of a campaign to characterize it in advance of a potential spacecraft encounter (e.g. [3]) which resulted in a rotation period of 3.77 h and an S-complex spectral classification (likely Sq or Q, [4,5,6]). The H magnitude was determined to be $H = 18.63\pm0.04$ and the R-band magnitude slope parameter to be $G = 0.18\pm0.02$ (formal errors; real errors up to 2x higher (P. Pravec, pers. comm.)). Potentially-hazardous asteroids (277475) 2005 WK₄ and (137126) 1999 CF₉ were observed with NSFCam2 on the IRTF [7] and were also successfully measured from the Goldstone and/or Arecibo radar observatories. Radar imaging of 2005 WK₄ shows a spheroidal shape and a diameter of $250\pm \approx 20$ m. SpeX data indicate that both 2005 WK₄ and 1999 CF₉ are also S-complex objects. Further analysis of these data is in progress.

> NSFCam2 Photometry of 2005 WK4, 2013-08-12 0.07 "2005 WK4 NSFCam2, 2013-08-12" 0.06 redicted Reflectance Based on V=13.98 Thermal Emission, Eta=1.0 0.05 Thermal Emission, Eta=1.5 Thermal Emission, Eta=1.75 Density (Jy) 0.04 Thermal Emission, Eta=2.0 **Ä** 0.03 0.02 0.014 0.01 0 0.7 1.2 1.7 2.2 2.7 3.2 5.2 3.7 4.7 Wavelength (Microns)

Figure: Preliminary thermal results from IRTF/NSFCam2 observations of 2005 WK₄.

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References: [1] A. T. Tokunaga, D. A. Simons, W. D. Vacca, PASP 114, 180 (2002). [2] K.W.Hodapp, et al., PASP 115,1388(2003). [3] J. Pittichova, et al., AAS/Division for Planetary Sciences. Meeting Abstracts (2013), vol. 45 of AAS/DPS, #101.06. [4] L. Franco, P. Bacci, L. Tesi, G. Fagioli, Minor Planet Bulletin 40, 205 (2013). [5] E. L. Ryan, et al., AAS/Division for Planetary Sciences Meeting Abstracts (2013), vol. 45 of AAS/DPS, #208.09. [6] F. Vilas, A. Hendrix, AAS/Division for Planetary Sciences Meeting Abstracts (2013), vol. 45 of AAS/DPS, #208.09. [6] F. Vilas, A. Hendrix, AAS/Division for Planetary Sciences Meeting Abstracts (2013), vol. 45 of AAS/DPS, #101.07. [7] M. A. Shure, D. W. Toomey, J. T. Rayner, P. M. Onaka, A. J. Denault, Instrumentation in Astronomy VIII, D. L. Crawford, E. R. Craine, eds. (1994), vol. 2198 of Society of Photo-Optical Instrumentation Engineers (SPIE) Conference Series, pp. 614–622. [8] J.R.Spencer, L.A.Lebofsky, M.V.Sykes, Icarus 78,337 (1989).