

Shape and thermal characterization of a candidate spacecraft target asteroid (175706) 1996 FG₃

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In this work, we have derived a 3D convex shape model for asteroid (175706) 1996 FG₃, which consists of 2040 triangle facets and 1022 vertices, from the known lightcurves. The best-fit orientation of the asteroid's spin axis is determined to be $\lambda = 237.7^\circ$ and $\beta = -83.8^\circ$, and its rotation period is ~ 3.5935 h. Using our shape model, we adopt the thermophysical model to fit three published sets of mid-infrared observations of 1996 FG₃ (Wolters 2011, Walsh 2012) to evaluate its surface properties. Assuming the primary and the secondary have identical shape, albedo, thermal inertia, and surface roughness, the best-fit parameters are obtained from the observations. The geometric albedo and effective diameter of the asteroid are calculated to be $p_v = 0.045 \pm 0.002$, $D_{\text{eff}} = 1.69_{-0.02}^{+0.05}$ km. The diameters of the primary and secondary are determined to be $D_1 = 1.63_{-0.03}^{+0.04}$ km and $D_2 = 0.45_{-0.03}^{+0.04}$ km, respectively. The surface thermal inertia Γ is derived to be a low value of $80 \pm 40 \text{ Jm}^{-2}\text{s}^{-0.5}\text{K}^{-1}$ with a roughness fraction f_R of $0.8_{-0.4}^{+0.2}$. The results suggest that the primary possibly has a regolith layer on its surface, which is likely to be covered by a mixture of dust, fragmentary rocky debris and sand. The minimum regolith depth is estimated to be $5 \sim 20$ mm from the simulations, indicating that 1996 FG₃ could be a very suitable candidate target for future space mission.

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